

5161

OXYGEN ISOTOPES IN MAFIC AND FELDSPATHIC CLASTS FROM POLYMICHT UREILITES

N. T. Kita¹, C. A. Goodrich², B. Fu¹, M. J. Spicuzza¹, and J. W. Valley¹.

¹Department of Geology and Geophysics, University of Wisconsin–Madison, Madison, Wisconsin, USA. E-mail: noriko@geology.wisc.edu. ²Department of Physical Sciences, Kingsborough Community College, Brooklyn, New York, USA

Introduction: Polymict ureilites DaG 165 and DaG 319 contain a wide variety of igneous clasts that might be derived from basaltic complements to the ultramafic monomict ureilites [1–2]. In a previous ion microprobe oxygen isotopic study of clasts from DaG 319, most plagioclase bearing clasts were found to plot along the CCAM-line and thus their origin from the ureilite parent body (UPB) was confirmed [1]. However, because of the small number of samples and limited precision of the analyses ($\delta^{17}\text{O} \sim 1\%$), oxygen isotopic systematics among these clasts is not well understood. In this study, the new ion microprobe IMS-1280 [3] was used to perform a systematic survey of 28 new clasts (described by [2, 4]) in DaG 165 and 319, including 20 plagioclase bearing clasts similar to those studied by [1]. In addition, two unusual mafic clasts (olivine with chromite exsolution, and Ferich pigeonite containing a melt inclusion) described by [4] and several of the olivine-augite feldspathic clasts described by [2] were studied to clarify their origin.

Ion Microprobe Analyses: The sample was sputtered with focused Cs⁺ primary ions (2–7 nA, 10–15 μm diameter) and oxygen three isotopes were analyzed by 3 Faraday Cup detectors combining mono and multicollection systems [3]. At least 4–6 repeated analyses were made for each clast in order to obtain $<0.3\%$ precisions in ^{18}O , ^{17}O , and ^{16}O . Olivine, pyroxene, and plagioclase standards with matching chemical compositions were used for correction of instrumental mass fractionation.

Results and Discussion: All but one of the feldspathic clasts plot along the CCAM line in the same region as ferroan monomict ureilites ($\delta^{17}\text{O} = \square 0.5$ to $\square 1.4\%$). One albitic plagioclase fragment plots on the TF line at $^{18}\text{O} \sim 6\%$, similar to enstatite meteorites. The two unusual mafic clasts plot on or slightly above the TF line at $^{18}\text{O} \sim 3\%$, indicating a possible link to primitive achondrites or ordinary chondrites. Four olivine-augite bearing feldspathic clasts plot along CCAM, confirming their UPB origin. None of the samples plot in the range of magnesian ureilites ($\delta^{17}\text{O} \sim \square 2\%$ on CCAM), consistent with the previous study [1]. If smelting determined mg#s of monomict ureilites and the UPB had a correlation of $\delta^{17}\text{O}$ with depth [4], our results indicate a bias toward materials from the deeper source regions (in contrast to the interpretation of [4] that the dominant feldspathic clasts were derived from the shallowest depths). This is likely due to mingling of melts during fractional melt extraction [6], a process that was not recognized in the modeling of [4]. Calcic clasts ($An > 30$) show slightly higher $\delta^{17}\text{O}$ than albitic clasts ($An < 30$) by 0.3%. Low degree melts enriched in alkali and other trace elements generated at depth might migrate into shallower region where isotopic mixing with lower $\delta^{17}\text{O}$ could occur. This is consistent with recent UPB models suggesting that Al-rich melt containing live ^{26}Al concentrated at shallower depths on the UPB [5, 6].

References: [1] Kita N. T. et al. 2004. *Geochimica et Cosmochimica Acta* 68:4213–4235. [2] Cohen B. A. et al. 2004. *Geochimica et Cosmochimica Acta* 68:4249–4266. [3] Kita N. T. 2006. Abstract #1496. 37th Lunar and Planetary Science Conference. [4] Goodrich C. A. et al. 2004. *Chemie de Erde* 64:283–327. [5] Kita N. T. et al. 2005. Abstract #5178. *Meteoritics & Planetary Science* 40. [6] Wilson et al. 2006. #1191. 37th Lunar and Planetary Science Conference.